

Describing limits of bounded sequences of measurable functions via nonstandard analysis

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XXVI incontro AILA - Padova

Different notions of limit - I

The Rademacher sequence

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- $\lim_{n \rightarrow \infty} r_n(x)$ does not exist for a.e. $x \in [0, 1]$;
- $r_n \rightarrow 0$ ($r_n \xrightarrow{*} 0$ if $p = 1, \infty$);

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- $\lim_{n \rightarrow \infty} r_n(x)$ does not exist for a.e. $x \in [0, 1]$;
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- for all $f \in C^0(\mathbb{R})$ and for all $\varphi \in C^0([0, 1])$,

$$\lim_{n \rightarrow \infty} \int_{[0,1]} f(r_n(x)) \varphi(x) dx = \left(\frac{1}{2} f(1) + \frac{1}{2} f(-1) \right) \int_{[0,1]} \varphi(x) dx.$$

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Definition (Young – 1937, ...)

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Theorem (The main theorem of Young measures)

For every bounded sequence of $L^p(\Omega)$ functions $\{z_n\}_{n \in \mathbb{N}}$, there exists a Young measure ν such that for all $f \in C_b^0(\mathbb{R})$ and for all $\varphi \in C^0([0, 1])$,

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$$z_n \xrightarrow{Y} \nu.$$



Nonstandard analysis in a nutshell

Let $\mathcal{P}(X)$ be the power set of X , and define

$$\mathbb{V}(X) = \bigcup_{n \in \mathbb{N}} \mathcal{P}^n(X).$$

Definition (Nonstandard universe)

A nonstandard universe is a triple $(\mathbb{V}(\mathbb{R}), \mathbb{V}({}^*\mathbb{R}), *)$ such that:

- $*$: $\mathbb{V}(\mathbb{R}) \rightarrow \mathbb{V}({}^*\mathbb{R})$;
- $*$ maps \mathbb{R} properly into ${}^*\mathbb{R}$ (i.e. $\mathbb{R} \neq {}^*\mathbb{R}$);
- $*$ preserves “elementary properties”.

The basic idea of nonstandard analysis

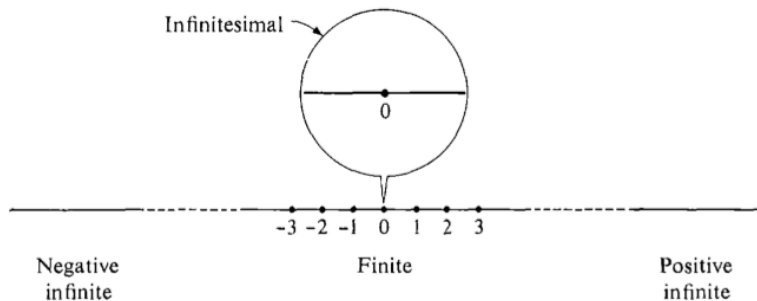


Image from "Elementary Calculus: An Infinitesimal Approach", © 2000 by H. Jerome Keisler.

Grid functions

Definition (Grid functions)

Let $N \in {}^*\mathbb{N}$ be infinite and let $\varepsilon = N^{-1}$. We define

$$\mathbb{X} = \{n\varepsilon : n \in {}^*\mathbb{Z} \text{ and } -N^2 \leq n \leq N^2\}.$$

The space of grid functions defined over an open set $\Omega \subseteq \mathbb{R}^k$ is

$$\mathbb{G}(\Omega) = \{f : {}^*\Omega \cap \mathbb{X}^k \rightarrow {}^*\mathbb{R}\}.$$

The main theorem

Theorem

For every bounded sequence $\{z_n\}_{n \in \mathbb{N}}$ in $L^p(\Omega)$, there exists a (non unique) function $z \in \mathbb{G}(\Omega)$ such that

- for all $\varphi \in C^0(\Omega)$ it holds

$$\varepsilon^k \sum_{x \in {}^*\Omega \cap \mathbb{X}^k} z(x)^* \varphi(x) \approx \lim_{n \rightarrow \infty} \int_{\Omega} z_n(x) \varphi(x) dx;$$

- for all $f \in C_b^0(\mathbb{R})$ and for all $\varphi \in C^0(\Omega)$ it holds

$$\varepsilon^k \sum_{x \in {}^*\Omega \cap \mathbb{X}^k} {}^*f(z(x))^* \varphi(x) \approx \lim_{n \rightarrow \infty} \int_{\Omega} f(z_n(x)) \varphi(x) dx.$$

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Theorem

For every distribution T , there exists $u \in \mathbb{G}(\Omega)$ that satisfies

$$\varepsilon^k \sum_{x \in {}^*\Omega \cap \mathbb{X}^k} u(x)^* \varphi(x) \approx \langle T, \varphi \rangle \text{ for all } \varphi \in C_c^\infty(\Omega).$$

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Theorem (Cutland – 1986)

For every Young measure ν , there exists $u \in \mathbb{G}(\Omega)$ that satisfies

$$\varepsilon^k \sum_{x \in {}^*\Omega \cap \mathbb{X}^k} {}^*f(u(x))^* \varphi(x) \approx \int_{\Omega} f(\nu(x)) \varphi(x) dx$$

for all $f \in C_b^0(\mathbb{R})$ and for all $\varphi \in C^0(\Omega)$

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- Let $z_n \rightarrow z_\infty$ ($z_n \xrightarrow{*} z_\infty$ if $p = 1, \infty$).

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- Let $z_n \xrightarrow{Y} \nu$, and define $b(x) = \int_{\mathbb{R}} \tau d\nu(x)$.

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$z = z_D + z_Y \in \mathbb{G}(\Omega)$ satisfies the desired properties.

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Perspectives for future research

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Possible applications of grid functions include:

- the definition of new notions of limits;
- the description of physical phenomena that cannot be formalized in the space of distributions;
- the study of generalized solutions to classically ill-posed problems from many areas of functional analysis.